

Magnetization relaxation and collective spin excitations in correlated double-exchange ferromagnets

M. D. Kapetanakis and I. E. Perakis

Department of Physics, University of Crete, Heraklion, 71003 Crete, Greece

and Institute of Electronic Structure and Laser, Foundation for Research and Technology–Hellas, Heraklion, 71110 Crete, Greece

(Received 31 May 2008; revised manuscript received 11 August 2008; published 10 October 2008)

We study spin relaxation and dynamics of collective spin excitations in correlated double-exchange ferromagnets. For this, we introduce an expansion of the Green's functions equations of motion that treats nonperturbatively all correlations between a given number of spin and charge excitations and becomes exact within a subspace of states. Our method treats relaxation beyond Fermi's golden rule while recovering previous variational results for the spin-wave dispersion. We find that the momentum dependence of the spin-wave dephasing rate changes qualitatively due to the on-site Coulomb interaction, in a way that resembles experiment, and depends on its interplay with the magnetic exchange interaction and itinerant spin lifetime. We show that the collective spin relaxation and its dependence on the carrier concentration depend sensitively on three-body correlations between a spin excitation and a Fermi sea electron and hole. The above spin dynamics can be controlled via the itinerant carrier population.

DOI: [10.1103/PhysRevB.78.155110](https://doi.org/10.1103/PhysRevB.78.155110)

PACS number(s): 75.30.Ds, 75.10.Lp, 75.47.Lx

I. INTRODUCTION

Long-range ferromagnetic order mediated by interactions between itinerant and localized spins is established in many different materials.^{1–4} The manganese oxides $R_{1-x}A_x\text{MnO}_3$ ($R=\text{La, Pr, Nd, Sm, \dots}$ and $A=\text{Ca, Ba, Sr, Pd, \dots}$) are prominent examples.² The magnetic and transport properties of such itinerant magnetic systems are intimately related and, unlike in other ferromagnets, can potentially be controlled by tuning parameters such as the itinerant carrier density.

Learning how to control magnetization dynamics and relaxation is important for spintronic applications.^{3,5,6} One of the challenges facing future magnetic devices and memories concerns their speed, which is governed by the dynamics of the collective spin. For small deviations from equilibrium, the dynamical magnetic properties are determined by the spin susceptibility.^{6–10} Within the random phase approximation (RPA),^{7,11–16} which gives the spin susceptibility to $O(1/S)$ (S is the local spin magnitude), magnetization relaxation arises from the interplay between the dephasing of the itinerant carrier spin and the magnetic exchange interaction.^{8–10,17} Spin relaxation also arises from inelastic-scattering processes such as magnon scattering with charge excitations.^{14,18,19} Experimental probes of such effects include neutron scattering, ferromagnetic resonance, and ultrafast magneto-optical pump-probe spectroscopy.^{17,20–27} The interpretation of such experiments requires the development of many-body theories of spin dynamics and relaxation.

Our goal in this paper is to develop a theory that describes the local spin Green's function

$$\langle\langle S_{\mathbf{Q}}^{\pm} \rangle\rangle = -i\theta(t)\langle[S_{\mathbf{Q}}^{\pm}(t), S_{\mathbf{Q}}^{\mp}(0)]\rangle, \quad (1)$$

which determines the transverse spin susceptibility. In the above we introduced the collective spin operators $\mathbf{S}_{\mathbf{q}}^n = 1/\sqrt{N}\sum_j \mathbf{S}_j^n e^{-i\mathbf{q}\cdot\mathbf{R}_j}$, $n=x, y, z$, where \mathbf{S}_j describe spins localized on N lattice sites at positions \mathbf{R}_j . $S^{\pm} = S^x \pm iS^y$ are the spin raising/lowering operators and $\langle\langle \dots \rangle\rangle$ denotes the grand canonical ensemble average. We consider a model Hamil-

tonian that accounts for the most important features common in a wide range of different itinerant ferromagnets $H=K+H_{\text{exch}}+H_{\text{AF}}+H_U$. $K=\sum_{\mathbf{k}\sigma}\epsilon_{\mathbf{k}}c_{\mathbf{k}\sigma}^{\dagger}c_{\mathbf{k}\sigma}$ describes a band of itinerant carriers, which in the manganites arises from the Mn d states with e_g symmetry. $c_{\mathbf{k}\sigma}^{\dagger}$ creates an electron with momentum \mathbf{k} , spin σ , and energy $\epsilon_{\mathbf{k}}$. To simplify the calculation of correlation effects common in many different physical systems, we consider a one-band model of $n=1-x$ itinerant electrons per Mn atom that hop between nearest-neighbor lattice sites. The electron concentration is described by the filling factor $n=N_e/N$, where N_e is the total number of electrons which varies from 0 to 1. $\epsilon_{\mathbf{k}}=-t\gamma_{\mathbf{k}}$, where $\gamma_{\mathbf{k}}=2\sum_{i=1}^d\cos(k_i a)$. d is the system dimensionality and a the lattice constant ($a=\hbar=1$ from now on). In the manganites, $0.5\leq n\leq 0.8$ in the metallic ferromagnetic regime of interest here. Our calculation can be extended to include the band structure of individual materials.

In momentum space, the magnetic exchange interaction between the local and itinerant spins is given by

$$H_{\text{exch}} = -\frac{J}{2\sqrt{N}}\sum_{\mathbf{k}, \mathbf{q}\sigma}\sigma S_{\mathbf{q}}^z c_{\mathbf{k}-\mathbf{q}\sigma}^{\dagger}c_{\mathbf{k}\sigma} - \frac{J}{2\sqrt{N}}\sum_{\mathbf{k}, \mathbf{q}}(S_{\mathbf{q}}^{\dagger}c_{\mathbf{k}-\mathbf{q}\uparrow}^{\dagger}c_{\mathbf{k}\downarrow} + \text{H.c.}), \quad (2)$$

where $\sigma=\pm 1$. In the manganites, Eq. (2) describes the Hund's rule onsite interaction between the e_g carrier spin and the $S=3/2$ local magnetic moment of the three electrons in the tightly bound t_{2g} Mn orbitals. $J\sim 2$ eV and 0.2 eV $\leq t \leq 0.5$ eV are typical values quoted in the literature, which give $4\leq J/t\leq 10$.²⁹ The Hamiltonian $K+H_{\text{exch}}$ defines the simple double-exchange model.^{28,29}

Here we add to the above minimal double exchange Hamiltonian two ubiquitous interactions;

$$H_U = \frac{U}{N} \sum_{\mathbf{k}\mathbf{k}'\mathbf{q}} c_{\mathbf{k}+\mathbf{q}\uparrow}^\dagger c_{\mathbf{k}'-\mathbf{q}\downarrow}^\dagger c_{\mathbf{k}'\downarrow} c_{\mathbf{k}\uparrow} \quad (3)$$

is the on-site (Hubbard) Coulomb repulsion. The Coulomb energy $U \sim 3.5-8$ eV is the largest energy scale in the manganites.²⁹ This Hubbard interaction is generally hard to treat and its effects on the spin dynamics have received less attention;^{14,15,30}

$$H_{AF} = J_{AF} \sum_{\mathbf{k}} \gamma_{\mathbf{k}} S_{\mathbf{k}}^z S_{-\mathbf{k}}^z + \frac{J_{AF}}{2} \sum_{\mathbf{k}} \gamma_{\mathbf{k}} (S_{\mathbf{k}}^+ S_{-\mathbf{k}}^- + S_{\mathbf{k}}^- S_{-\mathbf{k}}^+) \quad (4)$$

is the direct superexchange antiferromagnetic interaction between the nearest-neighbor local spins. $J_{AF} \sim 0.01t$ is weak in the manganites.²⁹

Given the large values of J/t in most systems of interest, many theories start from the strong-coupling limit ($J \rightarrow \infty$) of the above Hamiltonian. In this limit, the itinerant carriers can hop on a site only if their spin is parallel to the local spin there. The kinetic energy is reduced when all spins are parallel (double exchange mechanism²⁸), which favors the fully polarized half-metallic state $|F\rangle$. In the classical limit $S \rightarrow \infty$, the problem can be mapped to an effective nearest-neighbor Heisenberg model with ferromagnetic interactions. The lowest order [$O(1/S)$] quantum corrections are described by the RPA, which for strong couplings gives a dispersion that again coincides with that of the nearest-neighbor Heisenberg ferromagnet.¹¹ We may therefore assess the importance of correlations and quantum fluctuations beyond $O(1/S)$ by fitting the Heisenberg dispersion to the experimental result and looking for deviations.

For electron concentrations $n \geq 0.7$, initial measurements found nearest-neighbor Heisenberg model spin dynamics.³¹ However, later experiments reported deviations that increase strongly for $n \leq 0.7$.³²⁻³⁹ The Heisenberg model with nearest-neighbor interactions J_1 misses a pronounced softening near the Brillouin-zone boundary. This softening is accompanied by a strong increase in the spin-wave damping as we approach the zone boundary. The experimental dispersion could be fitted by adding a fourth-nearest-neighbor ferromagnetic exchange interaction J_4 to J_1 while keeping $J_3 = J_2 = 0$.³⁹

The above experimental observations reveal a spin dynamics and nonlocal correlations that are not captured by the strong-coupling limit of the double exchange model (which favors local correlations). Several scenarios have been put forward. The proposed mechanisms involve, among others, magnon scattering with orbital degrees of freedom,^{36,40} Fermi sea pairs^{30,41} and phonons,^{33,36,42} disorder effects,⁴³ band structure effects,⁴⁴ Hubbard interactions¹⁴ and correlations,³⁰ and the energetic overlap between spin-wave modes and the Stoner continuum.⁴⁵ The observed pronounced dependence of the spin-wave dynamics on the carrier concentration puts stringent conditions on the theory. Ye *et al.*³⁹ argued that none of the mechanisms proposed so far can fully account for all aspects of this spin dynamics.

The purpose of this paper is twofold. First, we study the momentum dependence of the spin-wave dephasing rate, with the focus on the role of correlations between the spin

and charge excitations and on their interplay with the itinerant spin dephasing. We demonstrate that both the magnitude and the momentum dependence of the spin relaxation rates depend sensitively on the carrier concentration and on correlations due to both J and U . We show that the on-site Coulomb repulsion U changes the momentum dependence of the spin-wave dephasing rate in a qualitative way that resembles the experimental results. We also show that our results depend on the interplay of U with the itinerant carrier spin lifetime, which is finite in some systems due to interactions not included in our Hamiltonian.^{7,8} We compare with the $1/S$ expansion and other approximations and find that three-body correlations between spin and electron-hole pair excitations play an important role. Finally, we show that the magnetization relaxation can be controlled by tuning the carrier density. We obtain changes in the spin relaxation with n that correlate with corresponding changes in the spin-wave softening and non-Heisenberg behavior.

Second, we develop and test a general method for describing spin dynamics. For this we use a truncation scheme of the infinite hierarchy of Green's function equations of motion based on an expansion in terms of correlations. Our scheme treats the full dynamics induced by the correlations between a given number of elementary excitations. Here we describe all correlations between a local or carrier spin excitation and an electron-hole Fermi sea pair and obtain the exact solution within the subspace of states with up to one Fermi sea excitation. Similar to Refs. 30 and 41, our method becomes exact in the limits of one electron ($N_e=1, n=1/N$), half filling ($N_e=N, n=1$), and in the atomic limit ($t=0$ for any n). It interpolates between the weak- and strong-coupling limits and agrees with exact diagonalization results for the spin-wave dispersion.⁴⁶ Finally, it retains its variational nature in the limit of zero relaxation rates, which provides a rigorous bound for the spin-wave softening and ferromagnetic phase boundary. Our approach, used before in the context of the Hubbard Hamiltonian,^{47,48} is in the same spirit as the projection and factorization scheme of Ref. 49, used to calculate the ultrafast nonlinear optical response of systems with a strongly correlated ground state and the correlation expansion of Ref. 50. It may be extended to study spin correlations in nonequilibrium systems^{17,27} and the magnetization dynamics of (III, Mn)V semiconductors.⁵¹

The rest of this paper is organized as follows. In Sec. II we discuss the Green's function truncation scheme and derive a closed system of equations of motion that determine the spin Green's function. In Sec. III we obtain the spin self-energy and separate the RPA contribution from the contributions of correlations due to J and U . In Sec. IV we discuss our numerical results for the spin-wave dephasing rate and dispersion and compare different approximations. In Sec. IV A we consider the minimal double exchange model with $U=J_{AF}=0$, while in Sec. IV B we study how U and J_{AF} change the picture. We end with our conclusions in Sec. V.

II. TRUNCATION OF GREEN'S FUNCTION HIERARCHY

In this section we obtain the equations of motion that determine the Green's function Eq. (1) and the spin suscep-

tibility. The many-body interactions H_{exch} and H_U introduce an infinite hierarchy of coupled equations of motion that involve higher Green's functions of the form

$$\langle\langle A \rangle\rangle = -i\theta(t)\langle[A(t), S_{\mathbf{Q}}^-(0)]\rangle, \quad (5)$$

where $A(t) = \exp(iHt)A \exp(-iHt)$ are many-body Heisenberg operators. To truncate this hierarchy, we approximate the higher Green's functions by systematically adding correlations among any given number of elementary excitations. To lowest order, RPA describes uncorrelated quasiparticles. At the next level, we include all correlations between any two elementary excitations, which determine the inelastic dephasing rate.

Reference 47 used a three-body scattering theory to calculate the electron Green's function of the Hubbard Hamiltonian. In one dimension, the results obtained this way were in excellent agreement with the exact Bethe ansatz solution.^{47,48} In Ref. 49, a similar method was used to calculate the density matrix that describes the coherent ultrafast nonlinear optical dynamics of the quantum hall system. References 52–54 calculated the Fermi edge singularity in doped semiconductor quantum wells using an analogous approach. In this paper, we establish the correspondence with a factorization scheme of higher Green's functions. For simplicity we restrict to zero temperature, where Eq. (5) involves the ground-state average value.

In the case of ferromagnetic exchange interaction as in the manganites, the fully polarized state

$$|F\rangle = \prod_{\nu} c_{\nu\uparrow}^{\dagger}|0\rangle \otimes |S, S, \dots\rangle \quad (6)$$

is an *exact eigenstate* of the many-body Hamiltonian H . In Eq. (6), $|0\rangle$ is the vacuum state and $|S, S, \dots\rangle$ describes local spins with $S_z = S$ on all lattice sites. From now on, the indices μ, ν, \dots denote states occupied in $|F\rangle$, while α, β, \dots denote empty states. In the parameter range of interest here, $|F\rangle$ is the ground state.^{30,41} Using the properties $H|F\rangle = 0$ (we choose the eigenvalue of $|F\rangle$ as the zero of energy) and $\langle F|S_{\mathbf{Q}} = 0$, both of which stem from the fact that $|F\rangle$ is the state with maximum spin, we obtain from Eq. (5)

$$\langle\langle A \rangle\rangle = -i\theta(t)\langle F|Ae^{-iHt}S_{\mathbf{Q}}|F\rangle. \quad (7)$$

The Green's function $\langle\langle A \rangle\rangle$ is then given by the amplitude of the time-evolved state $S_{\mathbf{Q}}|F\rangle$. The hierarchy of Green's function equations of motion is equivalent to solving the time-dependent Schrödinger equation. However, Green's functions also treat dephasing and relaxation, and can describe phenomenologically the effects of coupling to degrees of freedom not included in the Hamiltonian H by introducing phenomenological damping rates. The coupling of $\langle\langle A \rangle\rangle$ to higher Green's functions is determined by the states $HA|F\rangle$. Truncation of the equations of motion hierarchy can be achieved by expanding $HA|F\rangle$ in a truncated basis, which gives the exact solution within a subspace of states.

We start with the equation of motion for the spin Green's function Eq. (1), obtained after straightforward algebra by using Eq. (7) and the properties of $|F\rangle$:

$$\left(i\partial_t - \frac{Jn}{2} - \omega_{\mathbf{Q}}^{\text{AF}}\right)\langle\langle S_{-\mathbf{Q}}^{\dagger} \rangle\rangle = 2S\delta(t) - \frac{JS}{\sqrt{N}}\sum_{\nu}\langle\langle c_{\nu\uparrow}^{\dagger}c_{\nu+\mathbf{Q}\downarrow} \rangle\rangle + \frac{J}{2N}\sum_{\alpha\nu}\langle\langle S_{\alpha-\nu-\mathbf{Q}}^{\dagger}c_{\nu\uparrow}^{\dagger}c_{\alpha\uparrow} \rangle\rangle, \quad (8)$$

and

$$\omega_{\mathbf{Q}}^{\text{AF}} = 2J_{\text{AF}}S(\gamma_{\mathbf{Q}} - \gamma_0) \quad (9)$$

is the spin-wave energy due to H_{AF} . The same result can alternatively be obtained by decomposing the Green's functions contributing to $\langle\langle [S_{-\mathbf{Q}}^{\dagger}, H] \rangle\rangle$ into correlated and uncorrelated parts after using the identity

$$\begin{aligned} \langle\langle S_{\mathbf{q}}^n c_{\mathbf{k}-\mathbf{q}\sigma}^{\dagger} c_{\mathbf{k}\sigma'} \rangle\rangle &= \langle c_{\mathbf{k}-\mathbf{q}\sigma}^{\dagger} c_{\mathbf{k}\sigma'} \rangle \langle\langle S_{\mathbf{q}}^n \rangle\rangle \\ &+ \langle S_{\mathbf{q}}^n \rangle \langle\langle c_{\mathbf{k}-\mathbf{q}\sigma}^{\dagger} c_{\mathbf{k}\sigma'} \rangle\rangle \\ &+ \langle\langle \Delta S_{\mathbf{q}}^n \Delta [c_{\mathbf{k}-\mathbf{q}\sigma}^{\dagger} c_{\mathbf{k}\sigma'}] \rangle\rangle, \end{aligned} \quad (10)$$

where S^n are the components of the local spin and $\Delta A = A - \langle A \rangle$ describes the quantum fluctuations of A . The Green's function $\langle\langle \Delta S_{\mathbf{q}}^n \Delta [c_{\sigma}^{\dagger} c_{\sigma'}] \rangle\rangle$ is the correlated part of $\langle\langle S_{\mathbf{q}}^n c_{\sigma}^{\dagger} c_{\sigma'} \rangle\rangle$.

The Green's function $\langle\langle c_{\nu\uparrow}^{\dagger} c_{\nu+\mathbf{Q}\downarrow} \rangle\rangle$ describes the itinerant carrier spin dynamics and satisfies the following equation of motion, obtained after straightforward algebra by using Eq. (7) and the properties of $|F\rangle$:

$$\begin{aligned} &(i\partial_t - \varepsilon_{\nu+\mathbf{Q}} + \varepsilon_{\nu} - JS - nU)\langle\langle c_{\nu\uparrow}^{\dagger} c_{\nu+\mathbf{Q}\downarrow} \rangle\rangle \\ &= -\frac{J}{2\sqrt{N}}\langle\langle S_{-\mathbf{Q}}^{\dagger} \rangle\rangle - \frac{U}{N}\sum_{\mu}\langle\langle c_{\mu\uparrow}^{\dagger} c_{\mu+\mathbf{Q}\downarrow} \rangle\rangle \\ &- \frac{J}{2\sqrt{N}}\sum_{\alpha}\langle\langle S_{\alpha-\nu-\mathbf{Q}}^{\dagger} c_{\nu\uparrow}^{\dagger} c_{\alpha\uparrow} \rangle\rangle \\ &- \frac{U}{N}\sum_{\alpha\mu}\langle\langle c_{\mu\uparrow}^{\dagger} c_{\mathbf{Q}+\mu+\nu-\alpha\downarrow} c_{\nu\uparrow}^{\dagger} c_{\alpha\uparrow} \rangle\rangle_c, \end{aligned} \quad (11)$$

where we defined the correlated part of the four-particle Green's function as

$$\begin{aligned} \langle\langle c_1^{\dagger} c_2 c_3^{\dagger} c_4 \rangle\rangle_c &= \langle\langle c_1^{\dagger} c_2 c_3^{\dagger} c_4 \rangle\rangle - \langle c_1^{\dagger} c_2 \rangle \langle\langle c_3^{\dagger} c_4 \rangle\rangle - \langle c_3^{\dagger} c_4 \rangle \langle\langle c_1^{\dagger} c_2 \rangle\rangle \\ &- \langle c_2 c_3^{\dagger} \rangle \langle\langle c_1^{\dagger} c_4 \rangle\rangle + \langle c_1^{\dagger} c_4 \rangle \langle\langle c_3^{\dagger} c_2 \rangle\rangle. \end{aligned} \quad (12)$$

Alternatively, Eq. (11) can be derived by using Eq. (10) to decompose the Green's function $\langle\langle [c_{\nu\uparrow}^{\dagger} c_{\nu+\mathbf{Q}\downarrow}, H] \rangle\rangle$. The first line on the right-hand side (rhs) of Eqs. (8) and (11) gives the RPA result, which neglects all nonfactorizable (correlated) contributions to Eqs. (10) and (12) (Tyablikov approximation⁵⁵).

Equations (8) and (11) reduce the calculation of the spin Green's function to that of two higher Green's functions $\langle\langle \Delta S^{\dagger} \Delta [c_{\uparrow}^{\dagger} c_{\uparrow}] \rangle\rangle$ and $\langle\langle c_{\uparrow}^{\dagger} c_{\downarrow} c_{\uparrow}^{\dagger} c_{\downarrow} \rangle\rangle_c$. In a system with a general ground state, two additional Green's functions $\langle\langle \Delta S^z \Delta [c_{\uparrow}^{\dagger} c_{\downarrow}] \rangle\rangle$ and $\langle\langle \Delta S^{\dagger} \Delta [c_{\uparrow}^{\dagger} c_{\downarrow}] \rangle\rangle$ also couple and describe ground-state correlations.⁵¹ However, these vanish here since, for the ground state Eq. (6), $\Delta S^z|F\rangle = 0$ and $c_{\uparrow}^{\dagger} c_{\downarrow}|F\rangle = 0$. For the same reason, $\langle\langle c_{\mu\uparrow}^{\dagger} c_{\downarrow} c_{\nu\uparrow}^{\dagger} c_{\alpha\uparrow} \rangle\rangle = \langle\langle c_{\mu\uparrow}^{\dagger} c_{\downarrow} c_{\nu\uparrow}^{\dagger} c_{\alpha\uparrow} \rangle\rangle_c$ and $\langle\langle S^+ c_{\nu\uparrow}^{\dagger} c_{\alpha\uparrow} \rangle\rangle = \langle\langle \Delta S^+ \Delta [c_{\nu\uparrow}^{\dagger} c_{\alpha\uparrow}] \rangle\rangle$. Also, from Eq. (7) we see that $\langle\langle A \rangle\rangle = 0$ for any A such that $\langle F|A = 0$.

The Green's function $\langle\langle S^+ c_{\nu\uparrow}^\dagger c_{\alpha\uparrow} \rangle\rangle$ describes the correlations between a magnon, an electron, and a Fermi sea hole (three-body correlations), while the Green's function $\langle\langle c_{\mu\uparrow}^\dagger c_{\mathbf{Q}+\mu+\nu-\alpha\downarrow} c_{\nu\uparrow}^\dagger c_{\alpha\uparrow} \rangle\rangle_c$ describes the correlations between two Fermi sea holes and two electrons of opposite spin. These two Green's functions are obtained from the following equations of motion, derived from Eq. (7) after using the properties of $|F\rangle$:

$$\begin{aligned} & \left(i\partial_t - \varepsilon_\alpha + \varepsilon_\nu - \frac{Jn}{2} - \omega_{\mathbf{Q}+\nu-\alpha}^{\text{AF}} \right) \langle\langle S_{\alpha-\nu-\mathbf{Q}}^\dagger c_{\nu\uparrow}^\dagger c_{\alpha\uparrow} \rangle\rangle \\ &= \frac{J}{2N} \langle\langle S_{-\mathbf{Q}}^\dagger \rangle\rangle - \frac{JS}{\sqrt{N}} \langle\langle c_{\nu\uparrow}^\dagger c_{\nu+\mathbf{Q}\downarrow} \rangle\rangle \\ &+ \frac{J}{2N} \sum_\beta \langle\langle S_{\beta-\nu-\mathbf{Q}}^\dagger c_{\nu\uparrow}^\dagger c_{\beta\uparrow} \rangle\rangle \\ &- \frac{J}{2N} \sum_\mu \langle\langle S_{\alpha-\mu-\mathbf{Q}}^\dagger c_{\mu\uparrow}^\dagger c_{\alpha\uparrow} \rangle\rangle \\ &- \frac{JS}{\sqrt{N}} \sum_\mu \langle\langle c_{\mu\uparrow}^\dagger c_{\mathbf{Q}+\mu+\nu-\alpha\downarrow} c_{\nu\uparrow}^\dagger c_{\alpha\uparrow} \rangle\rangle_c \end{aligned} \quad (13)$$

and

$$\begin{aligned} & (i\partial_t - \varepsilon_{\mathbf{Q}+\mu+\nu-\alpha} - \varepsilon_\alpha + \varepsilon_\nu + \varepsilon_\mu - JS - nU) \\ & \times \langle\langle c_{\mu\uparrow}^\dagger c_{\mathbf{Q}+\mu+\nu-\alpha\downarrow} c_{\nu\uparrow}^\dagger c_{\alpha\uparrow} \rangle\rangle_c \\ &= \frac{J}{2\sqrt{N}} [\langle\langle S_{\alpha-\mu-\mathbf{Q}}^\dagger c_{\mu\uparrow}^\dagger c_{\alpha\uparrow} \rangle\rangle - \langle\langle S_{\alpha-\nu-\mathbf{Q}}^\dagger c_{\nu\uparrow}^\dagger c_{\alpha\uparrow} \rangle\rangle] \\ &+ \frac{U}{N} [\langle\langle c_{\mu\uparrow}^\dagger c_{\mu+\mathbf{Q}\downarrow} \rangle\rangle_c - \langle\langle c_{\nu\uparrow}^\dagger c_{\nu+\mathbf{Q}\downarrow} \rangle\rangle] \\ &- \frac{U}{N} \left[\sum_{\mu'} \langle\langle c_{\mu'\uparrow}^\dagger c_{\mathbf{Q}+\mu'+\nu-\alpha\downarrow} c_{\nu\uparrow}^\dagger c_{\alpha\uparrow} \rangle\rangle_c \right. \\ &+ \sum_{\nu'} \langle\langle c_{\mu\uparrow}^\dagger c_{\mathbf{Q}+\mu+\nu'-\alpha\downarrow} c_{\nu'\uparrow}^\dagger c_{\alpha\uparrow} \rangle\rangle_c \\ &\left. - \sum_{\alpha'} \langle\langle c_{\mu\uparrow}^\dagger c_{\mathbf{Q}+\mu+\nu-\alpha'\downarrow} c_{\nu\uparrow}^\dagger c_{\alpha'\uparrow} \rangle\rangle_c \right]. \end{aligned} \quad (14)$$

Equations (13) and (14) describe vertex corrections to the carrier-spin interaction. The first line on the rhs of Eq. (13) gives the Born approximation. The second and third lines describe vertex corrections due to the multiple scattering of the localized spin with the Fermi sea pair electron (second line) and hole (third line). Neglecting the third line corresponds to assuming a noninteracting (static⁵²) Fermi sea, equivalent to summing only the electron-magnon ladder diagrams (two-body ladder approximation).⁴⁷ By also including the hole multiple-scattering processes (third line), we treat exactly all correlations between local spin, electron, and hole; a three-body problem. The last line on the rhs of Eq. (13) comes from correlations between two electrons and two holes, described by Eq. (14). We note that U introduces new correlations among all four of the above particles, described by the last four lines on the rhs of Eq. (14).

To obtain the above closed system of equations, we neglected the coupling to Green's functions of the form $\langle\langle A c_{\nu\uparrow}^\dagger c_{\mu\uparrow}^\dagger c_{\alpha\uparrow} c_{\beta\uparrow} \rangle\rangle$, where $A=S^+$ or $c_{\nu'\uparrow}^\dagger c_{\downarrow}$. These neglected Green's functions describe multiparticle correlations between *two* Fermi sea pairs and a local spin or carrier spin-flip excitation which contribute to higher order in $1/S$. Alternatively, we can arrive at the same result by decomposing the Green's functions $\langle\langle S^+ c_{\uparrow}^\dagger c_{\uparrow}^\dagger c_{\uparrow} \rangle\rangle$ and $\langle\langle c_{\uparrow}^\dagger c_{\downarrow} c_{\uparrow}^\dagger c_{\uparrow} \rangle\rangle$ into uncorrelated and correlated parts, by separating out all possible factorizable contributions similar to Ref. 50, and neglecting the fully correlated contributions that describe correlations among three excitations. This correlation expansion neglects the contribution of states with two or more Fermi sea pair excitations and corresponds to an exact calculation of the Green's functions within the subspace of states with up to one Fermi sea pair. As discussed, e.g., in Refs. 47 and 48 and implied by Eq. (7), in the limit $\gamma, \Gamma \rightarrow 0$ the exact calculation of the Green's function within a given subspace is equivalent to the variational calculation of the spin-wave energy using a variational wave function that is a linear combination of the states that span the subspace (obtained in Refs. 30 and 41 for the problem at hand). One could include multipair correlations, e.g., by extending the approaches of Refs. 53, 54, and 56.

III. SPIN SELF-ENERGY

The spin self-energy can be calculated by solving the equations of motion derived above by Fourier transformation. Equation (8) gives

$$\langle\langle S_{-\mathbf{Q}}^\dagger \rangle\rangle_\omega = \frac{2S}{\omega - \omega_{\mathbf{Q}}^{\text{AF}} - \Sigma(\omega, \mathbf{Q})}, \quad (15)$$

where $\Sigma(\omega, \mathbf{Q})$ is the self-energy. Defining for convenience

$$X_\nu(\omega, \mathbf{Q}) = \frac{\langle\langle c_{\nu\uparrow}^\dagger c_{\nu+\mathbf{Q}\downarrow} \rangle\rangle_\omega}{\langle\langle S_{-\mathbf{Q}}^\dagger \rangle\rangle_\omega}, \quad (16)$$

$$G_{\alpha\nu}(\omega, \mathbf{Q}) = \frac{\langle\langle S_{\alpha-\nu-\mathbf{Q}}^\dagger c_{\nu\uparrow}^\dagger c_{\alpha\uparrow} \rangle\rangle_\omega}{\langle\langle S_{-\mathbf{Q}}^\dagger \rangle\rangle_\omega}, \quad (17)$$

$$\Phi_{\mu\nu}^\alpha(\omega, \mathbf{Q}) = \frac{\langle\langle c_{\mu\uparrow}^\dagger c_{\mathbf{Q}+\mu+\nu-\alpha\downarrow} c_{\nu\uparrow}^\dagger c_{\alpha\uparrow} \rangle\rangle_\omega}{\langle\langle S_{-\mathbf{Q}}^\dagger \rangle\rangle_\omega}, \quad (18)$$

and substituting into Eqs. (8), (11), (13), and (14) we express the self-energy in the form

$$\Sigma(\omega, \mathbf{Q}) = \frac{Jn}{2} - \frac{\Delta}{\sqrt{N}} \sum_\nu X_\nu + \frac{J}{2N} \sum_{\alpha\nu} G_{\alpha\nu}, \quad (19)$$

where $\Delta=JS$ is the magnetic energy. We calculate Σ nonperturbatively in the interactions and $1/S$ by solving the following coupled equations for X , G , and Φ :

$$\begin{aligned} & (nU + \Delta + \varepsilon_{\nu+\mathbf{Q}} - \varepsilon_\nu - \omega) X_\nu - \frac{U}{N} \sum_\mu X_\mu \\ &= \frac{J}{2\sqrt{N}} \left(1 + \sum_\alpha G_{\alpha\nu} \right) + \frac{U}{N} \sum_{\alpha\mu} \Phi_{\mu\nu}^\alpha, \end{aligned} \quad (20)$$

$$\begin{aligned} & \left(\omega + \varepsilon_\nu - \varepsilon_\alpha - \frac{Jn}{2} - \omega_{\mathbf{Q}-\alpha+\nu}^{\text{AF}} + i\gamma \right) G_{\alpha\nu} \\ &= \frac{J}{2N} \left(1 + \sum_\beta G_{\beta\nu} - \sum_\mu G_{\alpha\mu} \right) - \frac{\Delta}{\sqrt{N}} \left(X_\nu + \sum_\mu \Phi_{\mu\nu}^\alpha \right), \end{aligned} \quad (21)$$

where $\gamma \rightarrow 0$ and

$$\begin{aligned} & (nU + \Delta + \varepsilon_\alpha + \varepsilon_{\mathbf{Q}+\mu+\nu-\alpha} - \varepsilon_\mu - \varepsilon_\nu - \omega) \Phi_{\mu\nu}^\alpha \\ &= \frac{J}{2\sqrt{N}} (G_{\alpha\nu} - G_{\alpha\mu}) \\ &+ \frac{U}{N} (X_\nu - X_\mu) + \frac{U}{N} \left(\sum_{\mu'} \Phi_{\mu'\nu}^\alpha + \sum_{\nu'} \Phi_{\mu\nu'}^\alpha - \sum_\beta \Phi_{\mu\nu}^\beta \right). \end{aligned} \quad (22)$$

First we consider the RPA self-energy, obtained by setting $G = \Phi = 0$ in the above equations. We can then solve Eq. (20) analytically after noting that its solution has the form

$$X_\nu^{\text{RPA}}(\omega, \mathbf{Q}) = \frac{\chi^{\text{RPA}}(\omega, \mathbf{Q})}{nU + \Delta + \varepsilon_{\nu+\mathbf{Q}} - \varepsilon_\nu - \omega}. \quad (23)$$

Substituting the above expression into Eq. (20) we obtain

$$\chi^{\text{RPA}}(\omega, \mathbf{Q}) = \frac{J}{2\sqrt{N}} \frac{1}{1 - \frac{U}{N} \sum_\mu \frac{1}{nU + \Delta + \varepsilon_{\mu+\mathbf{Q}} - \varepsilon_\mu - \omega}}, \quad (24)$$

which gives after some straightforward algebra

$$X_\nu^{\text{RPA}}(\omega, \mathbf{Q}) = \frac{J}{2\sqrt{N}} \frac{1}{\Delta + U_{\nu\mathbf{Q}} + \varepsilon_{\nu+\mathbf{Q}} - \varepsilon_\nu - \omega}, \quad (25)$$

where we introduced the Coulomb-induced energy

$$U_{\nu\mathbf{Q}} = \frac{U}{N} \sum_\mu \frac{\varepsilon_{\mu+\mathbf{Q}} - \varepsilon_\mu - \varepsilon_{\nu+\mathbf{Q}} + \varepsilon_\nu}{nU + \Delta - i\Gamma + \varepsilon_{\mu+\mathbf{Q}} - \varepsilon_\mu - \omega}. \quad (26)$$

Substituting Eq. (25) into Eq. (19) after setting $G=0$ we obtain the RPA self-energy

$$\Sigma^{\text{RPA}}(\omega, \mathbf{Q}) = \frac{J}{2N} \sum_\nu \frac{U_{\nu\mathbf{Q}} + \varepsilon_{\nu+\mathbf{Q}} - \varepsilon_\nu - \omega}{\Delta - i\Gamma + U_{\nu\mathbf{Q}} + \varepsilon_{\nu+\mathbf{Q}} - \varepsilon_\nu - \omega}. \quad (27)$$

Following Ref. 7, in the above equations we added a phenomenological relaxation rate Γ that describes the itinerant spin lifetime due to interactions not included in our Hamiltonian H . This result can alternatively be obtained by substituting Δ by $\Delta - i\Gamma$, as derived with the Lindblad semigroup method in Ref. 17. In the intrinsic system described by the Hamiltonian H , $\Gamma \rightarrow 0$.

We now turn to the self-energy due to the correlations. By formally solving Eq. (20) for X_ν and substituting into Eq. (19), we separate the self-energy into RPA and correlated contributions $\Sigma(\omega, \mathbf{Q}) = \Sigma^{\text{RPA}}(\omega, \mathbf{Q}) + \Sigma^{\text{corr}}(\omega, \mathbf{Q})$. After some algebra we obtain that $\Sigma^{\text{corr}}(\omega, \mathbf{Q}) = \Sigma_J^{\text{corr}}(\omega, \mathbf{Q}) + \Sigma_U^{\text{corr}}(\omega, \mathbf{Q})$;

$$\Sigma_J^{\text{corr}} = \frac{J}{2N} \sum_{\alpha\nu} G_{\alpha\nu} \frac{U_{\nu\mathbf{Q}} + \varepsilon_{\nu+\mathbf{Q}} - \varepsilon_\nu - \omega}{\Delta - i\Gamma + U_{\nu\mathbf{Q}} + \varepsilon_{\nu+\mathbf{Q}} - \varepsilon_\nu - \omega} \quad (28)$$

is the contribution of the Fermi sea-magnon correlations due to J , described by the Green's function G , Eq. (21);

$$\Sigma_U^{\text{corr}} = -\frac{U}{N^{3/2}} \sum_{\alpha\nu\mu} \Phi_{\mu\nu}^\alpha \frac{\Delta - i\Gamma}{\Delta - i\Gamma + U_{\nu\mathbf{Q}} + \varepsilon_{\nu+\mathbf{Q}} - \varepsilon_\nu - \omega} \quad (29)$$

is the contribution of the Fermi sea pair-carrier spin-flip four-particle correlations described by Φ and discussed in Ref. 30. This latter contribution vanishes for $U=0$.

IV. NUMERICAL RESULTS

In this section we present our numerical results for the spin-wave dispersion, which we obtain by solving self-consistently the equation

$$\omega_{\mathbf{Q}} = \omega_{\mathbf{Q}}^{\text{AF}} + \text{Re } \Sigma(\omega_{\mathbf{Q}}, \mathbf{Q}) \quad (30)$$

and the spin-wave dephasing rate $\Gamma_{\mathbf{Q}}$, determined by $\text{Im } \Sigma(\omega_{\mathbf{Q}}, \mathbf{Q})$. We focus on the inelastic contribution to the dephasing rate due to the scattering between spin and charge excitations: $\Gamma_{\mathbf{Q}} = \Gamma_{\mathbf{Q}}^J + \Gamma_{\mathbf{Q}}^U$, where

$$\Gamma_{\mathbf{Q}}^J = -\text{Im } \Sigma_J^{\text{corr}}(\omega_{\mathbf{Q}}, \mathbf{Q}) \quad (31)$$

is the contribution due to the magnon-Fermi sea pair correlations described by G , Eq. (28), and

$$\Gamma_{\mathbf{Q}}^U = -\text{Im } \Sigma_U^{\text{corr}}(\omega_{\mathbf{Q}}, \mathbf{Q}) \quad (32)$$

is the contribution due to the carrier spin flip-Fermi sea pair correlations described by Φ , Eq. (22). An additional elastic contribution to the spin-wave lifetime can come from the imaginary part of the RPA self-energy, which however vanishes in the limit $\Gamma \rightarrow 0$ due to the finite carrier spin-flip excitation energy. Within the RPA, spin-wave dephasing can only arise from the interplay between the magnetic exchange interaction and carrier spin dephasing via external couplings.^{7,8}

Below we study the momentum dependence of $\Gamma_{\mathbf{Q}}$ along different directions in the Brillouin zone. We focus, in particular, on the directions $\Gamma-X$, $\Gamma-M$, and $X-M$, where $\Gamma = (0, 0)$, $X = (\pi, 0)$ and $M = (\pi, \pi)$. Our calculations were performed on a 20×20 square lattice, which as shown in Refs. 41 and 30 gives good convergence to the thermodynamic limit. Any small size effects are washed out when the relaxation rate γ in Eq. (21) exceeds the energy spacing.

A. Minimal double-exchange model

First we consider the simple double exchange Hamiltonian and set $U = J_{\text{AF}} = \Gamma = 0$. Figure 1 shows the spin-wave dispersion $\omega_{\mathbf{Q}}$ as function of the damping rate γ in Eq. (21). With decreasing γ , our results converge to the spin-wave dispersion obtained variationally in Ref. 41 and the $\gamma \rightarrow 0$ limit. Similar to the experiment,³⁹ our calculation can then be fitted to the dispersion of the Heisenberg Hamiltonian with first- and fourth-nearest-neighbor spin interactions J_1 and

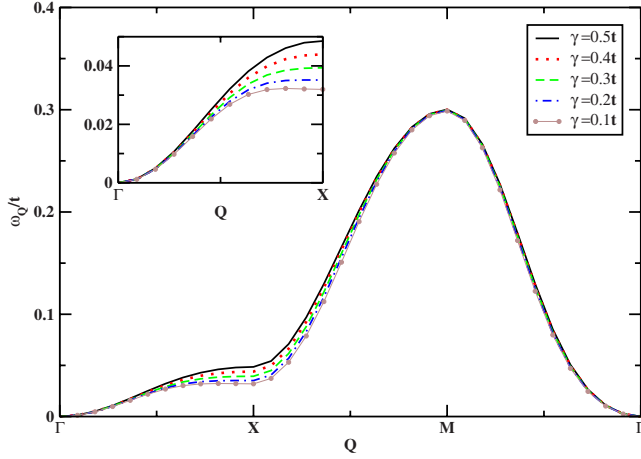


FIG. 1. (Color online) Spin-wave dispersion for different values of γ/t for $J=8t, n=0.6, U=J_{AF}=\Gamma=0$. Inset: Dispersion along the $\Gamma \rightarrow X$ direction.

J_4 .³⁰ For intermediate concentrations, our calculation gives a pronounced spin-wave softening at the X point, described by J_4 , as compared to both the RPA and to the fit to the nearest-neighbor Heisenberg dispersion. The effects of a finite γ are most pronounced along the direction $\Gamma \rightarrow X$: with increasing γ , the time evolution described by the Green's function G , which determines the spin-wave softening, is suppressed and thus the dispersion starts to approach the RPA ($G=0$) result. From now on we fix $\gamma=0.2t$, which as can be seen in Fig. 1 is close to the $\gamma \rightarrow 0$ limit.

Figure 2 demonstrates the important role of correlations due to spin-charge interactions on both the spin-wave dispersion and dephasing rate. The spin-wave energies and lifetimes differ markedly depending on the approximation used to treat the correlations. The latter determines the differences

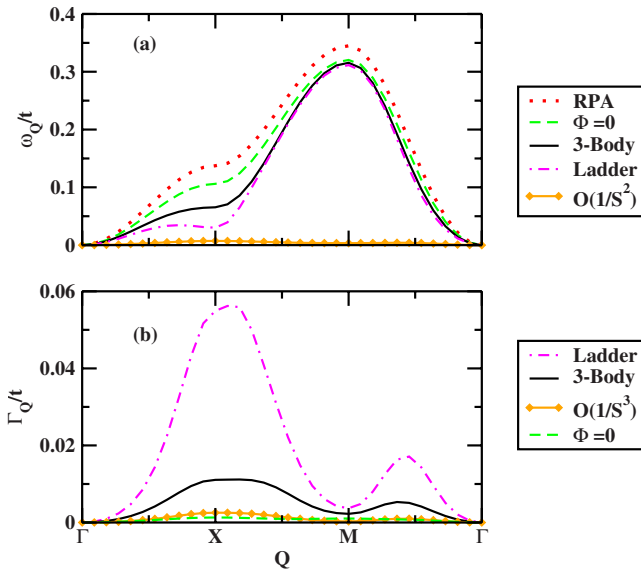


FIG. 2. (Color online) Comparison of the different approximations for treating the correlations for $U=0$. (a) Spin-wave dispersion and (b) inelastic spin-wave dephasing rate. $n=0.6, J=8t, \gamma=0.2t, U=J_{AF}=\Gamma=0$.

from the RPA, which describes noninteracting spin waves ($G=\Phi=0$). The RPA grossly underestimates the softening and does not give any spin damping in the limit $\Gamma \rightarrow 0$.

By neglecting the Green's function Φ , we obtain spin-wave energies closer to the RPA [see Fig. 2(a)]. As seen in Fig. 2(b), this approximation, which only treats the scattering of local spins with the Fermi sea, gives a very small damping rate. On the other hand, the nonvariational $O(1/S^2)$ approximation discussed in Refs. 14 and 19, which treats magnon-Fermi sea scattering within the Born approximation and Fermi's golden rule, strongly overestimates the softening, while at the same time predicting only a small damping rate (see Fig. 2).

By adding to the $O(1/S^2)$ result the effects of the multiple scattering of the magnon with the Fermi sea pair electron, while still neglecting the magnon-Fermi sea hole interactions, we obtain a nonvariational two-body approximation of the vertex corrections equivalent to summing the magnon-electron ladder diagrams.⁴⁷ As can be seen in Fig. 2, this ladder approximation gives very large softening *and* damping, much larger than the predictions of the full calculation (which is variational in the limit $\gamma, \Gamma \rightarrow 0$ considered here). The latter treats, in addition to the magnon-electron interactions, the multiple scattering of the Fermi sea pair hole with the magnon. The large differences between the ladder and full calculation results demonstrate the importance of three-body correlations between magnon, electron, and hole in the parameter regime of interest in the manganites. We conclude based on Fig. 2 that *all* correlations between spin, electron, and hole must be treated on an equal basis. The variational nature of our full calculation of the spin-wave energies in the limit $\gamma, \Gamma \rightarrow 0$ has the advantage of providing a rigorous limit of the magnitude of the softening, unlike for the ladder or $1/S$ expansion results.

Figure 3 shows the behaviors of ω_Q and Γ_Q for different interactions J . With increasing interaction strength, the spin-wave energies increase and the ferromagnetic phase becomes more stable. This hardening with J is accompanied by a corresponding increase in the spin-wave lifetime. The above changes are stronger along the Γ - X direction, where the non-Heisenberg behavior and softening are pronounced and depend nonlinearly on J/t . The overall momentum dependence, however, remains the same for all J .

To interpret the above results, we turn to the Green's function equations of motion and note that, for $U=0, \Sigma^U=0$. After solving Eqs. (20) and (22) for X and Φ and substituting into Eq. (21) we obtain that

$$\begin{aligned} \Omega_{\alpha\nu} G_{\alpha\nu} = & \frac{J}{2N} \frac{\varepsilon_{\nu+\mathbf{Q}} - \varepsilon_\nu - \omega}{\Delta - i\Gamma + \varepsilon_{\nu+\mathbf{Q}} - \varepsilon_\nu - \omega} \left(1 + \sum_\beta G_{\beta\nu} \right) \\ & - \frac{J}{2N} \sum_\mu G_{\alpha\mu} \\ & \times \frac{\varepsilon_{\mathbf{Q}+\mu+\nu-\alpha} - \varepsilon_\mu + \varepsilon_\alpha - \varepsilon_\nu - \omega}{\Delta - i\Gamma + \varepsilon_{\mathbf{Q}+\mu+\nu-\alpha} - \varepsilon_\mu + \varepsilon_\alpha - \varepsilon_\nu - \omega}, \end{aligned} \quad (33)$$

where

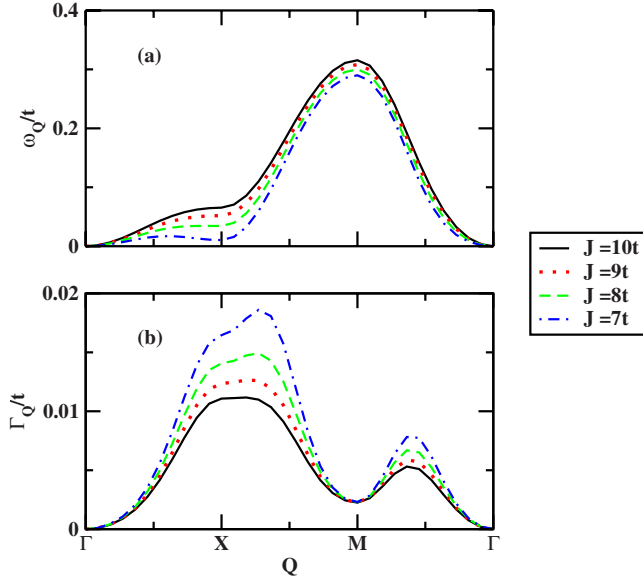


FIG. 3. (Color online) Dependence of the spin dynamics on the interaction strength J . (a) Spin-wave dispersion (b) and dephasing rate. $n=0.6, \gamma=0.2t, \Gamma=U=J_{AF}=0$.

$$\Omega_{\alpha\nu} = \omega + i\gamma - (\varepsilon_\alpha - \varepsilon_\nu) - \frac{J}{2N} \sum_{\mu} \frac{\varepsilon_{\mathbf{Q}+\mu+\nu-\alpha} - \varepsilon_{\mu} + \varepsilon_{\alpha} - \varepsilon_{\nu} - \omega}{\Delta - i\Gamma + \varepsilon_{\mathbf{Q}+\mu+\nu-\alpha} - \varepsilon_{\mu} + \varepsilon_{\alpha} - \varepsilon_{\nu} - \omega} \quad (34)$$

and $\gamma, \Gamma \rightarrow 0$.

Spin-wave dephasing results from the scattering of the magnon of momentum \mathbf{Q} to momentum $\mathbf{Q}+\nu-\alpha$ while an electron is excited from the state ν inside the Fermi sea to the empty state α . In the limit $\gamma \rightarrow 0$, this magnon-Fermi sea scattering process must satisfy the energy conservation condition $\Omega_{\alpha\nu}=0$, i.e., the initial magnon energy $\omega=\omega_{\mathbf{Q}}$ must equal the final-state energy that includes the Fermi sea pair energy and the $\mathbf{Q}+\nu-\alpha$ magnon energy. The final-state magnon energy, given by the last term in Eq. (34), comes from the coupling of G to Φ . For $\Phi=0$, this spin-wave energy is replaced by the local spin excitation energy $Jn/2$. The scattering of small energy Fermi sea pair excitations from right below to right above the Fermi surface dominates the spin-wave lifetime. The density of states and characteristic momenta of such pair excitations depend on the shape of the Fermi surface and therefore on the carrier concentration.

To derive the $O(1/S^3)$ dephasing rate,^{14,19} we neglect all rescattering terms on the rhs of Eq. (33) ($\propto G$). Substituting the expression for G obtained this way into Eq. (28), we obtain the Born-approximation self-energy

$$\Sigma_J^B = \frac{J^2}{4N^2} \sum_{\alpha\nu} \left(\frac{\varepsilon_{\nu+\mathbf{Q}} - \varepsilon_{\nu} - \omega}{\Delta + \varepsilon_{\nu+\mathbf{Q}} - \varepsilon_{\nu} - \omega} \right)^2 \frac{1}{\Omega_{\alpha\nu}}. \quad (35)$$

Expanding in terms of $1/S$ while keeping $\Delta=JS$ fixed and substituting $\omega=\omega_{\mathbf{Q}}^{(1)}$, where $\omega_{\mathbf{Q}}^{(1)}$ denotes the $O(1/S)$ spin-

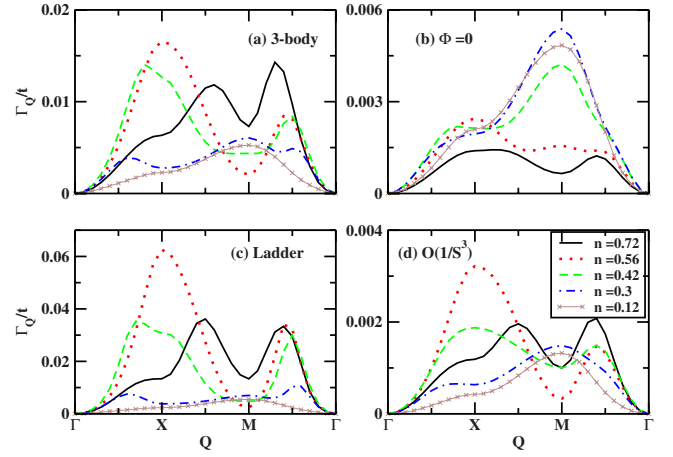


FIG. 4. (Color online) Dependence of the spin-wave dephasing rate on the carrier concentration n for $U=0$: comparison between the different approximations of the correlations $J=8t, \gamma=0.2t, \Gamma=U=J_{AF}=0$.

wave energy, we obtain the lowest-order contribution to the self-energy imaginary part

$$\text{Im} \Sigma_J^B(\omega_{\mathbf{Q}}, \mathbf{Q}) \approx \frac{\Delta^2}{4N^2 S^2} \sum_{\alpha\nu} \left(\frac{\varepsilon_{\nu} - \varepsilon_{\nu+\mathbf{Q}}}{\varepsilon_{\nu} - \varepsilon_{\nu+\mathbf{Q}} - \Delta} \right)^2 \text{Im} \frac{1}{\Omega_{\alpha\nu}}, \quad (36)$$

where

$$\Omega_{\alpha\nu} = \omega_{\mathbf{Q}}^{(1)} - (\varepsilon_{\alpha} - \varepsilon_{\nu} + \omega_{\mathbf{Q}-\alpha+\nu}^{(1)}) - i\gamma. \quad (37)$$

The above result corresponds to the Fermi's golden rule description of the magnon lifetime. Its large difference from our full calculation, demonstrated by Fig. 2, is due to the magnon-electron and magnon-hole multiple interactions (vertex corrections), described by the terms proportional to G on the rhs of Eq. (33). The comparison between the different approximations shows that, in the parameter regime of interest in the manganites, the vertex corrections due to three-body correlations renormalize significantly the magnon-carrier scattering.

We finally turn to the dependence of the spin relaxation on the carrier concentration. Figure 4(a) demonstrates a strong n -dependence of $\Gamma_{\mathbf{Q}}$, which correlates with an analogous dependence of $\omega_{\mathbf{Q}}$ and the spin-wave softening discussed in Ref. 30. As n decreases and the softening (non-Heisenberg behavior) disappears, the spin-wave lifetime increases while its momentum dependence changes. For intermediate n , $\Gamma_{\mathbf{Q}}$ displays two sharp peaks and a dip as function of momentum. For small n , the overall $\Gamma_{\mathbf{Q}}$ decreases and the positions of its maxima and minima change.

To see this concentration dependence in more detail, we note that, for $n=0.72$, the spin-wave damping is maximized for $\mathbf{Q} \sim (\pi, \pi/2)$, between X and M , and $\mathbf{Q} \sim (\pi, \pi/2)$, between M and Γ , while it is minimized close to the M point.

As n decreases to intermediate values, the first of the above maxima approaches the X point while the second maximum shifts closer to $\mathbf{Q}=(\pi, \pi/2)$. For smaller n , the dip close to the M point turns into a maximum. As a result of this n dependence, $\Gamma_{\mathbf{Q}}/\omega_{\mathbf{Q}}$ becomes quite large in the direction $\Gamma \rightarrow X$ for intermediate n , which implies that these spin-wave quasiparticles interact strongly with the Fermi sea. For small n , $\Gamma_{\mathbf{Q}}/\omega_{\mathbf{Q}}$ decreases again. The changes in the momentum dependence with n are related to the changes in the shape and position of the Fermi surface, which is located close to the Brillouin-zone boundary for the higher n but moves toward the center of the Brillouin zone as n decreases. As a result, the phase space available for magnon-carrier scattering changes drastically with n .

Figure 4 also compares the concentration dependence of $\Gamma_{\mathbf{Q}}$ predicted by the different approximations of the spin-charge interactions. By comparing the full three-body calculation with the $O(1/S^3)$ Fermi's golden rule result, it is clear that the spin-wave damping is grossly underestimated by the perturbative $1/S$ expansion for all concentrations. Figure 4(b), obtained by setting $\Phi=0$, fails completely to capture the correct concentration dependence [compare Figs. 4(a) and 4(b)]. It also predicts very small dephasing rates for all n . The above approximation neglects the interactions between Fermi sea pair and carrier spin-flip excitations. We therefore conclude that such carrier-carrier interactions strongly affect the magnetization relaxation. Finally, the comparison of Figs. 4(a) and 4(c) shows that the two-body ladder approximation grossly overestimates the spin-wave damping for intermediate or high n , while the discrepancies from the full three-body calculation decrease for small n . We conclude based on Fig. 4 that the collective spin relaxation predicted by the minimal double-exchange model can be controlled by tuning the carrier concentration n , by doping, or with external probes. Such tuning is heavily influenced by the correlations, which must be treated accurately in order to capture even the correct order of magnitude and momentum dependence of the spin dephasing rate for all concentrations.

B. The role of the Coulomb repulsion

In this section we study how the on-site Coulomb repulsion U and direct superexchange interaction J_{AF} affect the spin-wave energies and lifetimes. Figure 5 compares the results obtained for different values of J_{AF}/t within the range $0 \leq J_{AF} \leq 0.012t$ relevant to the manganites.²⁹ J_{AF} leads to an overall softening of the spin-wave energies. These eventually turn negative, implying instability of the fully polarized ferromagnetic phase. However, J_{AF} preserves the nearest-neighbor Heisenberg model momentum dependence, unlike for the softening observed in the experiment.³⁹ Also, it does not affect the spin damping in a significant way. We take $J_{AF}=0$ from now on.

As demonstrated by Fig. 6, the effects of the on-site Coulomb repulsion (Hubbard) repulsion U are more significant. Figure 6(a) shows the dependence of the spin-wave dispersion on U . With increasing U , the spin-wave softening and deviations from the Heisenberg model dispersion diminish as the

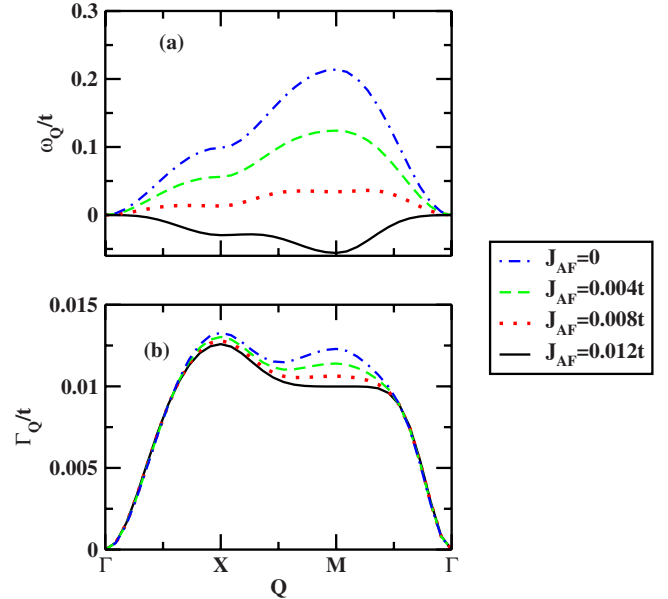


FIG. 5. (Color online) Role of the superexchange interaction J_{AF} . (a) Spin-wave dispersion and (b) spin-wave damping rate. $J=2t, U=10t, n=0.6, \Gamma=0.5t, \gamma=0.2t$.

ferromagnetic phase becomes more stable. As discussed in Ref. 30, our calculated dispersion can be fitted to the Heisenberg model dispersion with first- and fourth-nearest-neighbor interactions, similar to the experiment.³⁹ Figure 6(b) demonstrates qualitative changes in the overall momentum dependence of the spin-wave damping as compared to the minimal double exchange model. In particular, if the carrier spin is conserved ($\Gamma=0$), $\Gamma_{\mathbf{Q}}$ is maximum at the M point, while the damping at the X point is smaller. In contrast, for $U=0$ and intermediate concentrations, the dephasing rate displays a dip at the M point and is maximum close to the X point (see, e.g., Fig. 3). The double-peak momentum dependence of $\Gamma_{\mathbf{Q}}$ for $U=0$ can be recovered for large U only by introducing a sufficiently large itinerant spin damping rate Γ [see Fig. 6(c)]. The experiment of Ref. 39 observed an increase in the

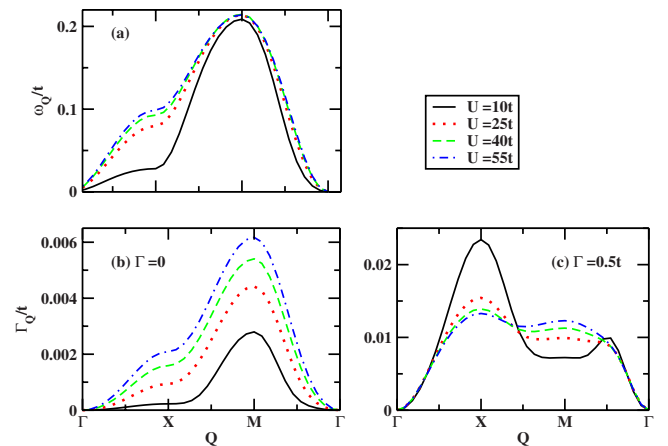


FIG. 6. (Color online) Effects of the on-site Coulomb repulsion U . (a) Spin-wave dispersion, (b) spin-wave damping rate for $\Gamma=0$, and (c) spin-wave damping rate for $\Gamma=0.5t$. $n=0.6, J=2t, \gamma=0.2t, J_{AF}=0$.

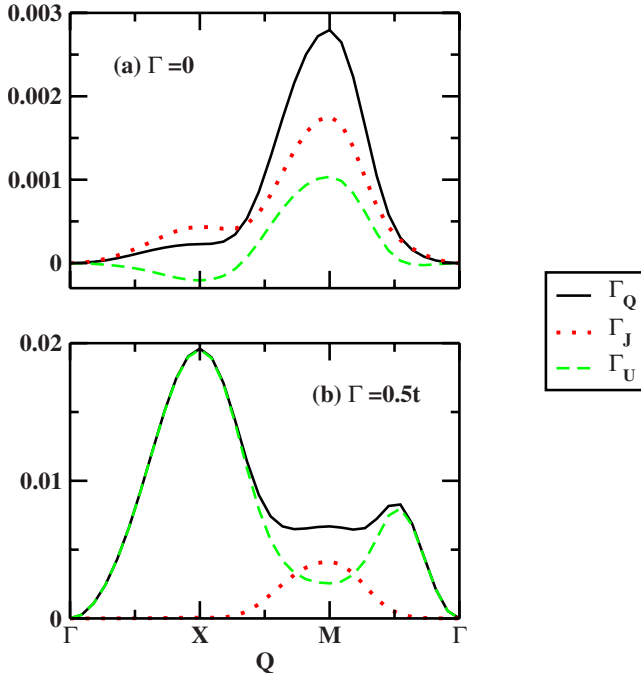


FIG. 7. (Color online) Contributions Γ_Q^J and Γ_Q^U to the inelastic spin-wave dephasing rate Γ_Q . (a) $\Gamma=0$ and (b) $\Gamma=0.5t$. $n=0.6$, $J=2t$, $U=10t$, $\gamma=0.2t$, $J_{AF}=0$.

spin-wave damping along $\Gamma \rightarrow M$ that exceeds the corresponding increase along $\Gamma \rightarrow X$, similar to our results for finite U and $\Gamma=0$ [Fig. 6(b)]. Our calculations show that such behavior of the spin damping can be attributed to the Coulomb repulsion U in the intrinsic system ($\Gamma, \gamma \rightarrow 0$) described by the Hamiltonian H .

Figure 6 also demonstrates a qualitative difference in the dependence of Γ_Q on U between the cases of small and large itinerant spin dampings. For $\Gamma=0$, the spin-wave dephasing rate *increases* with U [Fig. 6(b)], while for $\Gamma=0.5t$ it *decreases* with U close to the X point [Fig. 6(c)]. This result indicates that the magnetization relaxation may depend on the interplay between Coulomb repulsion and the dephasing of the itinerant spin via the coupling to an external bath.

To interpret the above behaviors, we plot in Fig. 7 the two contributions to Γ_Q obtained from the self-energies Eqs. (28) and (29) for zero and finite Γ . Γ^J is determined by G , while Γ^U is determined by Φ . For $\Gamma=0$, Fig. 7(a) shows that Γ^J and Γ^U are comparable in magnitude, since in this case they both arise from $\text{Im } G$. On the other hand, as Γ increases, the relative magnitude of Γ^J and Γ^U changes and the latter dominates [see Fig. 7(b)]. This enhancement of Γ^U arises from the additional contribution to $\text{Im } \Phi$, Eq. (22), obtained by adding the relaxation rate Γ to Eq. (22). Even though Γ^J continues to have the same momentum dependence as for $\Gamma=0$, Γ^U does not [compare Figs. 7(a) and 7(b)].

Figure 8 shows the dependences of the two contributions to the inelastic dephasing rate on the Coulomb repulsion for large U and different Γ . Γ^J increases and then saturates with increasing U . For $\Gamma=0$ (intrinsic system), Γ^U increases with U and eventually exceeds Γ^J . Unlike for Γ^J , the dependence of Γ^U on the momentum and on U is qualitatively different for large and small Γ . For example, Γ^U decreases with U at

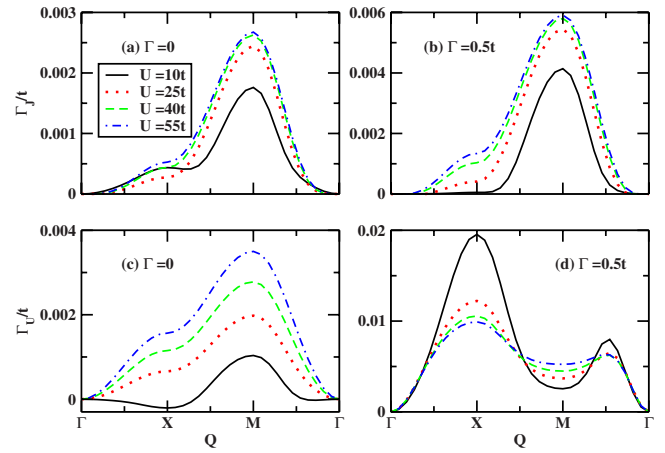


FIG. 8. (Color online) Dependences of Γ_Q^J and Γ_Q^U on Coulomb repulsion for large U . $n=0.6$, $J=2t$, $\gamma=0.2t$, $J_{AF}=0$.

the X point for large Γ but increases for $\Gamma=0$. The above behavior of Γ^U dominates the total dephasing rate for large Γ .

Figure 9(a) shows the transition in the momentum dependence of Γ_Q as U increases for $\Gamma=0$. This transition occurs around $U \sim 6t$, where the double peak momentum dependence for $U=0$, with a dip at the M point, changes into a peak at the M point. As can be seen in Fig. 9(b), the above transition arises from the changes in the behavior of Γ^J , determined by the Green's function G [Eq. (21)] that are introduced by the Coulomb repulsion. Figure 9(c), on the other hand, shows that the overall momentum dependence of Γ^U remains approximately the same for all U .

Finally, we turn to the possibility of controlling the magnetization relaxation by tuning the carrier concentration n and study how the Hubbard repulsion U changes the picture as compared to the prediction of Fig. 4. Figure 10 shows the dependence of the dispersion and spin damping rate on n within a wide range of concentrations $n=0.7-0.1$. As can be seen in Fig. 10(a), the pronounced softening along the Γ - X direction disappears rapidly with decreasing n (or increasing

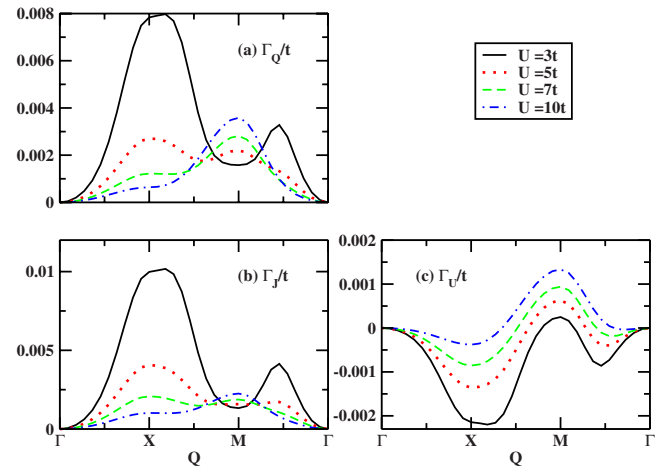


FIG. 9. (Color online) Dependences of (a) Γ_Q , (b) Γ_Q^J , and (c) Γ_Q^U on the Coulomb repulsion for small U . $n=0.6$, $J=4t$, $\gamma=0.2t$, $J_{AF}=\Gamma=0$.

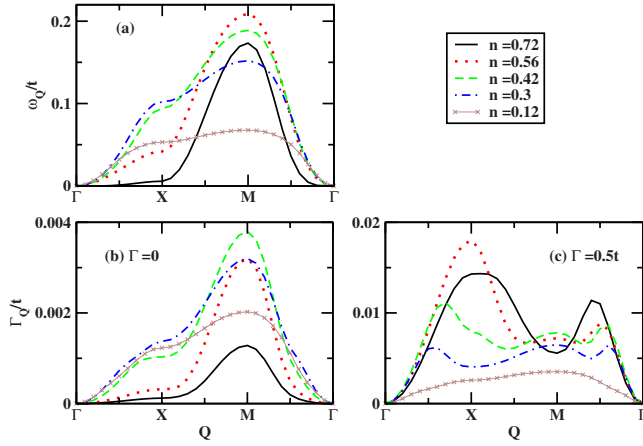


FIG. 10. (Color online) Dependence of spin-wave dispersion and dephasing rate on the carrier concentration n for $U=10t$. (a) Dispersion, (b) dephasing rate for $\Gamma=0$, and (c) dephasing rate for $\Gamma=0.5t$. $J=2t, J_{AF}=0, \gamma=0.2t$.

hole doping $x=1-n$). For small values of n , the overall energies decrease and the overall shape of the dispersion changes.

The above n dependence of the spin-wave dispersion correlates with corresponding changes in the dephasing rate. Figure 10(b), obtained for $\Gamma=0$, shows that the Coulomb repulsion U changes drastically the dependence of Γ_Q on n as compared to the predictions of the minimal double exchange model, Fig. 4(a). In this case, the spin dephasing rate *increases* as n decreases to intermediate values, in a way correlated with the disappearance of the spin-wave softening and non-Heisenberg behavior. For small n , where the softening has disappeared and the overall energies start to decrease, the spin damping rate also decreases. Furthermore, unlike for $U=0$ or for large Γ , Γ_Q displays a strong *maximum* at point M for *all concentrations* and is always weaker along the Γ - X direction. As can be seen by comparing Figs. 10(a) and 10(b), a sufficiently short itinerant spin lifetime (large Γ) changes drastically the momentum dependence of Γ_Q and its dependence on n .

We conclude based on Fig. 10 that the magnitude and momentum dependence of Γ_Q , as well as the spin-wave softening, can be controlled by tuning the carrier concentration n via hole doping or by external means such as photoexcitation.^{17,20-27} It is clear that the on-site Coulomb (Hubbard) repulsion plays a dominant role by inducing new correlations and dynamics absent in the simple double-exchange model. Such correlations change drastically the momentum dependence and magnitude of Γ_Q with varying n and must be treated in a consistent way in order to arrive at trustworthy conclusions and comparisons to experiment. Our results suggest that, as a first step, a systematic experimental study of the magnetization dynamics as function of doping x and a comparison to the theory are necessary in order to decide which many-body mechanisms dominate the collective magnetization dynamics and relaxation and learn how to control this dynamics for potential magnetic device and spintronics applications.

V. CONCLUSIONS

In this paper we presented a general method for describing the spin-wave dynamics and relaxation in itinerant ferromagnets. This method is based on a correlation expansion of the Green's function equations of motion that systematically treats all correlations between any given numbers of elementary excitations. Using this method, we derived a closed system of equations that treats the magnetic exchange and Coulomb interactions nonperturbatively and solved it to obtain the Green's function that determines the transverse spin susceptibility. Our results for the spin-wave dispersion reproduce previous variational^{30,41} and exact diagonalization^{46,47} results (in the limit $\Gamma, \gamma \rightarrow 0$) and therefore allow us to draw definite conclusions regarding the magnitude of the spin-wave softening. Using the properties of the fully polarized Hartree-Fock ground state with maximum spin, we showed that our method gives the exact spin Green's function within a subspace of states that includes up to one Fermi sea pair excitation. Our factorization scheme of the higher Green's functions also applies to other ground states. Our results recover the $1/S$ expansion results as special case. We showed that, in the parameter regime of interest in the manganites, the latter approximation overestimates the spin-wave softening, while at the same time it grossly underestimates the spin-wave damping rate. Furthermore, by comparing with the ladder approximation treatment of the vertex corrections to the magnon-carrier scattering, which treats the multiple magnon-electron scatterings while neglecting the interactions with a Fermi sea hole, we showed that three-body correlations between the magnon and an electron-hole Fermi sea pair excitation have an important effect on the spin relaxation.

Using the above many-body theory, we calculated the inelastic spin-wave dephasing rate nonperturbatively in the interactions and $1/S$ (i.e., beyond the standard Fermi's golden rule). We showed that correlations between a carrier spin-flip excitation and a Fermi sea pair induced by the Coulomb repulsion U play a very important role in the parameter regime relevant to the manganites. We also showed that both the magnitude and momentum dependence of the spin-wave dephasing rate depend sensitively on the itinerant carrier concentration. This result implies the possibility of controlling the magnetization relaxation in itinerant ferromagnets by tuning the carrier concentration, either via doping or by external means (e.g., photoexcitation or by using electric fields and currents or gates). We also argued that the interplay between on-site Coulomb (Hubbard) interaction and a finite itinerant carrier spin lifetime, due to the coupling of the different spin states induced by spin orbit or other interactions not included in our Hamiltonian, can affect our results in the realistic system. The momentum dependence of the spin-wave dephasing rate observed in recent experiments³⁹ is consistent with the results of our calculation only for sufficiently large U and $\Gamma \rightarrow 0$. In all other cases, we obtain a distinctly different double peak momentum dependence. Although the band structure of the relevant materials must be included in order to arrive at quantitative comparisons with the experiment, our calculation already demonstrates the crucial role of correlations. The agreement of the main trends in spin-wave

softening and damping rate changes as function of n between our theory and the experiment suggests that the simple one-band model already contains the main inelastic-scattering processes and correlations. Our results suggest that ultrafast magneto-optical pump-probe spectroscopy experiments, which directly probe the changes in spin relaxation and dynamics induced by photoexcited carriers,¹⁷ may provide

insight into the physics of the manganites and other itinerant ferromagnetic systems.

ACKNOWLEDGMENT

This work was supported by the EU STREP program HYSWITCH.

- ¹E. L. Nagaev, Phys. Rep. **346**, 387 (2001).
- ²See e.g., *Colossal Magnetoresistive Oxides*, edited by Y. Tokura (Gordon Breach, Singapore, 2000), and references therein.
- ³See e.g., Y. Tokura, Phys. Today **56** (7), 50 (2003).
- ⁴T. Jungwirth, J. Sinova, J. Masek, J. Kucera, and A. H. MacDonald, Rev. Mod. Phys. **78**, 809 (2006).
- ⁵S. A. Wolf, D. D. Awschalom, R. A. Buhrman, J. M. Daugherton, S. von Molnar, M. L. Roukes, A. Y. Chtchelkanova, and D. M. Treger, Science **294**, 1488 (2001).
- ⁶Y. Tserkovnyak, A. Brataas, G. Bauer, and B. I. Halperin, Rev. Mod. Phys. **77**, 1375 (2005).
- ⁷B. Heinrich, D. Fraitova, and V. Kambersky, Phys. Status Solidi **23**, 501 (1967).
- ⁸A. H. Mitchell, Phys. Rev. **105**, 1439 (1957).
- ⁹J. Sinova, T. Jungwirth, X. Liu, Y. Sasaki, J. K. Furdyna, W. A. Atkinson, and A. H. MacDonald, Phys. Rev. B **69**, 085209 (2004).
- ¹⁰Y. Tserkovnyak, G. A. Fiete, and B. I. Halperin, Appl. Phys. Lett. **84**, 5234 (2004).
- ¹¹N. Furukawa, J. Phys. Soc. Jpn. **65**, 1174 (1996).
- ¹²X. Wang, Phys. Rev. B **57**, 7427 (1998).
- ¹³E. L. Nagaev, Phys. Rev. B **58**, 827 (1998).
- ¹⁴D. I. Golosov, Phys. Rev. Lett. **84**, 3974 (2000); Phys. Rev. B **71**, 014428 (2005).
- ¹⁵S.-J. Sun, W.-C. Lu, and H. Chou, Physica B (Amsterdam) **324**, 286 (2002).
- ¹⁶M. Marder, N. Papanicolaou, and G. C. Psaltakis, Phys. Rev. B **41**, 6920 (1990); L. R. Mead and N. Papanicolaou, *ibid.* **28**, 1633 (1983).
- ¹⁷J. Chovan, E. G. Kavousanaki, and I. E. Perakis, Phys. Rev. Lett. **96**, 057402 (2006); J. Chovan and I. E. Perakis, Phys. Rev. B **77**, 085321 (2008); J. Chovan, E. G. Kavousanaki, and I. E. Perakis, Phys. Status Solidi C **3**, 2410 (2006).
- ¹⁸V. Yu. Irkhin and M. I. Katsnelson, Eur. Phys. J. B **19**, 401 (2001), and references therein.
- ¹⁹N. Shannon and A. V. Chubukov, Phys. Rev. B **65**, 104418 (2002); J. Phys.: Condens. Matter **14**, L235 (2002).
- ²⁰M. Vomir, L. H. F. Andrade, L. Guidoni, E. Beaurepaire, and J.-Y. Bigot, Phys. Rev. Lett. **94**, 237601 (2005).
- ²¹M. van Kampen, C. Jozsa, J. T. Kohlhepp, P. LeClair, L. Lagae, W. J. M. de Jonge, and B. Koopmans, Phys. Rev. Lett. **88**, 227201 (2002).
- ²²A. V. Kimel, A. Kirilyuk, A. Tsvetkov, R. V. Pisarev, and Th. Rasing, Nature (London) **429**, 850 (2004).
- ²³J. Qi, Y. Xu, N. H. Tolk, X. Liu, J. K. Furdyna, and I. E. Perakis, Appl. Phys. Lett. **91**, 112506 (2007).
- ²⁴D. Talbayev, H. Zhao, G. Lüpke, A. Venimadhav, and Qi Li, Phys. Rev. B **73**, 014417 (2006).
- ²⁵A. V. Kimel, A. Kirilyuk, F. Hansteen, R. V. Pisarev, and Th. Rasing, J. Phys.: Condens. Matter **19**, 043201 (2007).
- ²⁶J. Wang, C. Sun, Y. Hashimoto, J. Kono, G. A. Khodaparast, L. Cywinski, L. J. Sham, G. D. Sanders, C. Stanton, and H. Munekata, J. Phys.: Condens. Matter **18**, R501 (2006).
- ²⁷T. V. Shahbazyan, I. E. Perakis, and M. E. Raikh, Phys. Rev. Lett. **84**, 5896 (2000).
- ²⁸C. S. Zener, Phys. Rev. **82**, 403 (1951); P. W. Anderson and H. Hasegawa, *ibid.* **100**, 675 (1955); K. Kubo and N. Ohata, J. Phys. Soc. Jpn. **33**, 21 (1972).
- ²⁹E. Dagotto, T. Hotta, and A. Moreo, Phys. Rep. **344**, 1 (2001).
- ³⁰M. D. Kapetanakis and I. E. Perakis, Phys. Rev. B **75**, 140401(R) (2007).
- ³¹T. G. Perring, G. Aeppli, S. M. Hayden, S. A. Carter, J. P. Remmeika, and S.-W. Cheong, Phys. Rev. Lett. **77**, 711 (1996).
- ³²H. Y. Hwang, P. Dai, S.-W. Cheong, G. Aeppli, D. A. Tennant, and H. A. Mook, Phys. Rev. Lett. **80**, 1316 (1998).
- ³³P. Dai, H. Y. Hwang, J. Zhang, J. A. Fernandez-Baca, S.-W. Cheong, C. Kloc, Y. Tomioka, and Y. Tokura, Phys. Rev. B **61**, 9553 (2000).
- ³⁴L. Vasiliu-Doloc, J. W. Lynn, A. H. Moudden, A. M. de Leon-Guevara, and A. Revcolevschi, Phys. Rev. B **58**, 14913 (1998).
- ³⁵T. Chatterji, L. P. Regnault, and W. Schmidt, Phys. Rev. B **66**, 214408 (2002).
- ³⁶Y. Endoh, H. Hiraka, Y. Tomioka, Y. Tokura, N. Nagaosa, and T. Fujiwara, Phys. Rev. Lett. **94**, 017206 (2005).
- ³⁷T. Chatterji, L. P. Regnault, P. Thalmeier, R. Suryanarayanan, G. Dhalenne, and A. Revcolevschi, Phys. Rev. B **60**, R6965 (1999).
- ³⁸N. Shannon, T. Chatterji, F. Ouchni, and P. Thalmeier, Eur. Phys. J. B **27**, 287 (2002).
- ³⁹F. Ye, P. Dai, J. A. Fernandez-Baca, H. Sha, J. W. Lynn, H. Kawano-Furukawa, Y. Tomioka, Y. Tokura, and J. Zhang, Phys. Rev. Lett. **96**, 047204 (2006); F. Ye, P. Dai, J. A. Fernandez-Baca, D. T. Adroja, T. G. Perring, Y. Tomioka, and Y. Tokura, Phys. Rev. B **75**, 144408 (2007).
- ⁴⁰G. Khaliullin and R. Kilian, Phys. Rev. B **61**, 3494 (2000).
- ⁴¹M. D. Kapetanakis, A. Manousaki, and I. E. Perakis, Phys. Rev. B **73**, 174424 (2006).
- ⁴²D. M. Edwards, Adv. Phys. **51**, 1259 (2002).
- ⁴³Y. Motome and N. Furukawa, Phys. Rev. B **71**, 014446 (2005).
- ⁴⁴I. V. Solov'yev and K. Terakura, Phys. Rev. Lett. **82**, 2959 (1999).
- ⁴⁵T. A. Kaplan, S. D. Mahanti, and Y.-S. Su, Phys. Rev. Lett. **86**, 3634 (2001).
- ⁴⁶J. Zang, H. Röder, A. R. Bishop, and S. A. Trugman, J. Phys.: Condens. Matter **9**, L157 (1997); T. A. Kaplan and S. D. Mahanti, *ibid.* **9**, L291 (1997).

- ⁴⁷J. Igarashi, J. Phys. Soc. Jpn. **52**, 2827 (1983); **54**, 260 (1985); J. Igarashi, M. Takahashi, and T. Nagao, *ibid.* **68**, 3682 (1999).
- ⁴⁸A. E. Ruckenstein and S. Schmitt-Rink, Int. J. Mod. Phys. B **3**, 1809 (1989).
- ⁴⁹I. E. Perakis and E. G. Kavousanaki, Chem. Phys. **318**, 118 (2005); A. T. Karathanos, I. E. Perakis, N. A. Fromer, and D. S. Chemla, Phys. Rev. B **67**, 035316 (2003); K. M. Dani, E. G. Kavousanaki, J. Tignon, D. S. Chemla, and I. E. Perakis, Solid State Commun. **140**, 72 (2006).
- ⁵⁰J. Fricke, Ann. Phys. (N.Y.) **252**, 479 (1996); J. Fricke, V. Meden, C. Wohler, and K. Schönhammer, *ibid.* **253**, 177 (1997); V. Meden, C. Wohler, J. Fricke, and K. Schönhammer, Phys. Rev. B **52**, 5624 (1995).
- ⁵¹M. D. Kapetanakis and I. E. Perakis, Phys. Rev. Lett. **101**, 097201 (2008).
- ⁵²J. F. Mueller, A. E. Ruckenstein, and S. Schmitt-Rink, Phys. Rev. B **45**, 8902 (1992); A. E. Ruckenstein and S. Schmitt-Rink, *ibid.* **35**, 7551 (1987).
- ⁵³I. E. Perakis and Y.-C. Chang, Phys. Rev. B **47**, 6573 (1993); **44**, 5877 (1991).
- ⁵⁴T. V. Shahbazyan, N. Primožich, I. E. Perakis, and D. S. Chemla, Phys. Rev. Lett. **84**, 2006 (2000); N. Primožich, T. V. Shahbazyan, I. E. Perakis, and D. S. Chemla, Phys. Rev. B **61**, 2041 (2000).
- ⁵⁵S. V. Tyablikov, *Methods in the Quantum Theory of Magnetism* (Plenum, New York, 1967).
- ⁵⁶W. von der Linden and D. M. Edwards, J. Phys.: Condens. Matter **3**, 4917 (1991).